

#### Neutrinos were "invented" in 1930 by W.Pauli to explain some features of nuclear beta decay

Dear Radioactive Ladies and Gentlemen,

As the bearer of these lines, to whom I graciously ask you to listen, will explain to you in more detail, how because of the "wrong" statistics of the N and Li6 nuclei and the continuous beta spectrum, I have hit upon a desperate remedy to save the "exchange theorem" of statistics and the law of conservation of energy. Namely, the possibility that there could exist in the nuclei electrically neutral particles, that I wish to call *neutrons*, which have spin 1/2 and obey the exclusion principle and which further differ from light guanta in that they do not travel with the velocity of light. The mass of the *neutrons* should be of the same order of magnitude as the electron mass and in any event not larger than 0.01 proton masses. The continuous beta spectrum would then become understandable by the assumption that in beta decay a neutron is emitted in addition to the electron such that the sum of the energies of the neutron and the electron is constant......

I agree that my remedy could seem incredible because one should have seen those neutrons very earlier if they

Marinar - Ploto and of all 0393 Absohrist/15.12.5 Öffener Brief an die Gruppe der Radioaktiven bei der Genversing-Tagung an Tibingen. Absohrift Physikelisches Institut der Eidg. Technischen Hochschule Zirich, 4. Des. 1930 Wrich. Oloriastrasse Liebe Radioaktive Damen und Herren, Wie der Veberbringer dieser Zeilen, den ich huldvollat ansuhören bitte, Ihnen des näheren auseinendersetsen wird, bin ich angesichts der "felschen" Statistik der N- und Li-6 Kerne, sowie des kontinuierlichen bete-Spektruns auf einen versweifelten Ausweg verfallen um den "Wecheelsets" (1) der Statistik und den Energiesats su retten. Mamlich die Möglichkeit, es könnten elektrisch neutrals Telloben, die ich Neutronen nennen will, in den Iernen existieren, weiche dem Spin 1/2 heben und das Ausschliessungsprinzip befolgen und

aten von Lichtquanten musserden noch dadurch unterwoheiden, dass sie nicht mit Lichtquanten musserden noch dadurch unterwoheiden, dass sie Namete von derzalben Grossenordnung wis die Elektronenmasse sein und jehnsfalls nicht grösser als 0.00. Protonenmasses.- Das kontinuierliche Bein-Spektrum wire dann verständlich unter der Annahme, dass bein bein-Zerfall mit dem Elektron jeweils noch ein Neutron emittiert migel, derart, dass die Summe der Energien von Meutron und Elektron konsten tist.

really exist. But only the one who dare can win and the difficult situation, due to the continuous structure of the beta spectrum, is lighted by a remark of my honoured predecessor, Mr Debye, who told me recently in Bruxelles: "Oh, It's well better not to think to this at all, like new taxes". From now on, every solution to the issue must be discussed. Thus, dear radioactive people, look and judge. Unfortunately, I cannot appear in Tubingen personally since I am indispensable here in Zurich because of a ball on the night of 6/7 December. With my best regards to you, and also to Mr Back.

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Energy and momentum conservation fix the energy in 2-body decays Example:  $N \rightarrow N'+a$ 





The continuum of energies observed in  $\beta$  decay implies that either energy and momentum are not conserved ...or that there is another particle difficult to detect

 $N \rightarrow N' + \beta + v$ 

*Neutrinos* 

## Today neutrinos occupy an important role in particle physics...



#### So... where can we find neutrinos ?

## A Google search for "neutrino" gives about 7,390,000 hits !

#### ...a clear sign that neutrinos are everywhere !

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#### Sources of neutrinos: artificial and natural



#### We know of 4 types of fundamental interactions in nature:

- Gravity: every object is affected
  - (it is actually a property of space-time)
- Weak force: affects most particles including neutrinos
- Electromagnetism: affects all particles with charge
- Strong force: affect only quarks

Neutrinos having no electric charge interact only by the weak force (and gravity) <u>→ interact very little</u>

Strength

#### **Radioprotection for neutrinos ?**

 Mean free path of neutrinos from a reactor in lead is ~ 0.3 light years !



•A big nuclear reactor makes 6.10<sup>20</sup> neutrinos/s: at 20 meter distance (just outside the building) only one neutrino every 3 sec interacts with our body !



Bethe & Peierls 1934: "... this implies that one evidently TRIUMF never will be able to detect Neutrinos."





Detectors tend to be large and require extensive shielding against natural radioactivity

Material where neutrinos interact and are detected (1000 tons liquid scintillator... ever bought liquid scintillator in... bulk ?)

Shielding oil (2.5 m thick, or about 2 kton)

2000 photomultipliers (20 inch diameter)

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#### Make your own scintillator, underground



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#### And it even works !







## This is the optical analogous of the sonic boom produced by an aircraft







### Let's use these tools ...











#### **Conclusion 1:**

We detect vs ! nuclear fusion powers the sun

The sun is still shining (this is not trivial:
it takes ~ 1Myr for a photon to emerge from the sun)



to the neutrinos in their journey from the core of the sun to the earth





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Data collected with the SNO detector: • 1 kton of heavy water • ~2000 m underground in Ontario • World of Neutrinos

In Quantum Mechanics there are 2 representations for our neutrinos if  $m_v \neq 0$ :

So... what happens ?

 "Weak interaction eigenstate"



this is the state V<sub>e</sub> V<sub>μ</sub> of definite flavor: interactions couple to this state

"Mass eigenstate"



 $\begin{pmatrix} V_{m1} \\ V_{m2} \end{pmatrix}$  this is mean and a second state the second state state the state state

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#### The 2 eigenstates are connected by a $3 \cdot 3$ matrix ("mixing matrix")

$$\begin{pmatrix} v_{e} \\ v_{\mu} \\ v_{\tau} \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix} \begin{pmatrix} v_{m1} \\ v_{m2} \\ v_{m3} \end{pmatrix}$$

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A source produces -say-  $v_e$   $V_e$  always via weak interactions e

To see what propagates to the detector I have to project this flavor state onto the mass state using the matrix inverse of U

 $v_m = U^{-1} v_e$ 

Now each of the  $v_m$ will evolve in time as prescribed by the wave functions

$$V_{m1}(t) = e^{-i(E_1t - p_1L)}V_{m1}$$
$$V_{m2}(t) = e^{-i(E_2t - p_2L)}V_{m2}$$
$$V_{m3}(t) = e^{-i(E_3t - p_3L)}V_{m3}$$
note that the periodic term  
tains the neutrino mass via  
E:=m;c<sup>2</sup>

So at the end of their flight -at the detectorthe mix of  $v_{m1}$ ,  $v_{m2}$ ,  $v_{m3}$  will not necessarily be the one that makes "exactly"  $v_e!$ 

but

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The neutrinos are then detected via weak interactions and so we need to find again the composition in terms of  $v_e$ ,  $v_\mu$ ,  $v_\tau$ 

#### Formally

$$\left|\nu_{j}\right\rangle = \sum_{j'} \sum_{l} U_{lj} e^{-i(E_{j}t - p_{j}L)} U_{j'l}^{*} \left|\nu_{j'}\right\rangle$$

So after some propagation one can "find" a neutrino of a flavor that was not originally present

This is a pure quantum-mechanical effect, there is no classical interpretation of it (this is why it looks so weird)

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The situation is analogous to the case of beats in sound

The wavefunctions of the two neutrino mass eigenstates need to be combined in a linear combination to obtain the flux of a particular flavor eigenstate

The beat represents the oscillating intensity of a detected neutrino flavor beats\beats.html

The mass (together with the kinetic energy) of the neutrino give the frequency of the wavefunction

So this quantum mechanical "beat" phenomenon allows us to detect extremely small neutrino masses

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#### Can we detect oscillations in phenomena other than solar neutrinos ?

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69 reactor cores from Japan and (South) Korea

Baseline is limited: 85.3% of signal has 140 km < L < 344 km

The total electric power produced "as a by-product" of the vs is: ~60 GW or... ~4% of the world's manmade power or... ~20% of the world's nuclear power

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#### The result from reactors agrees very well with what expected from the phenomenon of solar neutrinos



The World of Neutrinos

So, we have now observed neutrino oscillations and hence we know that neutrinos have a finite mass

However measuring beats only provide information about the *difference* between two frequencies...

 $\rightarrow$  We still do not know the neutrino mass scale !

#### Summary of knowledge on neutrino masses



Allright, this is lots of fun but... can we use these neutrinos for something useful ?!

The answer is usually "no", among other things because neutrinos are so difficult to detect...

... but, remember the story of M. Faraday and the King...

Let's go back to a device that makes LOTS of antineutrinos: a nuclear reactor We can turn things around and use antineutrinos to "peek inside" the reactor's core (neutrinos don't care that there are heavy walls !)

#### The antineutrino count rate varies in a known way as Pu is produced even at constant power



Technology demonstration detector being prepared at the San Onofre Nuclear Generating Station (CA) [Sandia Natl Lab]

Set the 1m<sup>3</sup> detector in the "tendon gallery" -24.5 m from the core

Heavy reactor building provides shielding from cosmic rays

Expect 2600 v/day with 40% detector efficiency





## Earth sciences and the neutrinos Structure

From seismic data 5 basic regions:

of the

Earth

- inner core,
- outer core,
- mantle,
- oceanic crust,
- continental crust and sediments
- All these regions behave like solids except the outer core.

#### Only a shallow layer has been sampled for chemical composition by drilling/sampling



Global distribution of heat flow data. 38,347 data points from Davies and Davies, 2010.

- Deepest bore-hole (12km) is only ~1/500 of the Earth's radius
- Oceans and southern hemisphere substantially less studied

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The world of medulinos





#### Supernovae and the neutrinos





"type II" SN in the Crab;

Jul 4T,héOCrabA Nebula formed by the (as expelled duternshell of the same star Anasazi in Cha(asCreçorded/by the VLT, and the ancienCernonPar)anal, Chile)

#### Type II supernovae:

explosive phase of a star with M>6 to 8  $M_{sun}$ 

- Nuclear fuel burnt through Fe: no mechanism to hold further gravitational collapse
- T=0.8\*10<sup>10</sup> K = 0.7 MeV  $\rho=3*10^9 g/cm^3$  (this is a billion times the density of the earth, or the entire KamLAND in a teaspoon)
- The pressure causes the reaction  $p+e \rightarrow n+v_e$  to occur

Very intense v<sub>e</sub> flash (~1 s duration)

- Neutrinos cool the star that collapses further
- The collapsing soup become so dense to be opaque to v(!)
- Following mechanisms that we do not completely understand the fireball re-bounces blowing up the outer envelope (eventually like the Crab photo)

• v of all flavors escape when density low enough (after ~10 s) TRIUMF Jan 18, 2013

#### SN1987A reconstruction from <u>real</u> astronomical images





Only hours later the density is low enough for light to escape the star and the supernova flash appears

#### We have seen the neutrinos from a supernova only once

Feb 23, 1987 about 20 v events were detected from an explosion in the Large Magellanic Cloud: SN1987A



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The World of Neutrinos



Relic neutrinos have been "red-shifted" by the by the expansion of the Universe

Their temperature now is supposed to be ~2K, lower than the temperature (3K) of the microwave background radiation that was produced when the Universe was 30,000 years old

We do not have the technology to directly detect these neutrinos

But... we can see how their presence shaped the density fluctuations of matter in the early Universe as imaged by the microwave

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# Neutrinos,

mysterious ambassadors from the microworld to the remote cosmos, what lessons do you have for us next?